Exact and asymptotic formulas for overdamped Brownian dynamics

P.J. Forrester¹, B. Jancovici*

Laboratoire de Physique Théorique et Hautes Energies², Université de Paris-Sud,
91405 Orsay Cedex, France

Received 16 October 1996

Abstract

Exact and asymptotic formulas relating to dynamical correlations for over-damped Brownian motion are obtained. These formulas include a generalization of the $f$-sum rule from the theory of quantum fluids, a formula relating the static current−current correlation to the static density−density correlation, and an asymptotic formula for the small-$k$ behaviour of the dynamical structure factor. Known exact evaluations of the dynamical density−density correlation in some special models are used to illustrate and test the formulas.

PACS: 05.20.-y; 05.40.+j

1. Introduction

A number of recent papers [1--4] have considered various aspects of the density and current correlations in the Dyson Brownian motion model [5]. This model refers to the overdamped Brownian dynamics of the one-dimensional one-component log-gas. When the initial state is given by the equilibrium state, the model is equivalent [1] to the ground state dynamics of the quantum many body system with $1/r^2$ two-body interactions (Calogero–Sutherland model). One motivating factor for the current interest is that the Dyson Brownian motion model specifies the eigenvalue probability density function (p.d.f.) for certain ensembles of parameter dependent random matrices (see e.g. [6]). These random matrix ensembles are used to study the parametric dependence (typically as a function of magnetic field strength) of quantum energy spectra and fluctuation phenomena in quantum transport problems [7,8]. Another motivating factor...

* E-mail: janco@stat.th.u-psud.fr.
¹ Permanent address: Department of Mathematics, University of Melbourne, Parkville, Victoria 3052, Australia. E-mail: matpj@maths.mu.oz.au.
² Laboratoire associé au Centre National de la Recherche Scientifique. URA D0063.
is that there are now a number of exact results available for the dynamical density-density correlation \([4, 7-9]\).

In this paper we seek to extend the considerations of earlier works by considering properties of dynamical correlations for overdamped Brownian motion described by the general Fokker–Planck equation

\[
\gamma \frac{\partial p}{\partial \tau} = \mathcal{L} p \quad \text{where} \quad \mathcal{L} = \sum_{j=1}^{N} \frac{\partial}{\partial x_j} \left( \frac{\partial W}{\partial x_j} + \beta^{-1} \frac{\partial}{\partial x_j} \right).
\]

(1.1)

The Dyson Brownian motion model is specified by this equation with the particular potential

\[
W = \frac{1}{2} \sum_{j=1}^{N} x_j^2 - \sum_{1 \leq j < k \leq N} \log |x_k - x_j|, \quad (1.2a)
\]

or its periodic version

\[
W = - \sum_{1 \leq j < k \leq N} \log |e^{2\pi i x_j / L} - e^{2\pi i x_k / L}|. \quad (1.2b)
\]

We begin Section 2 by reviewing properties of (1.1). We then clarify the meaning of the current (notice that (1.1) is independent of the velocities). Next a generalization of the \(f\)-sum rule for quantum fluids \([10]\) is given, as is a generalization of the formula of Taniguchi et al. \([3]\) relating the initial density-density and current-current correlations. In the final two subsections of Section 2 we consider the hydrodynamic approximation to the microscopic density fluctuations, and derive from it the small-\(k\) asymptotics of the density-density and current-current correlations, which in turn imply certain sum rules and further asymptotic formulas. The analytic formulas of Section 2 are illustrated and tested on some exact results for the density-density correlation in the Dyson Brownian motion model (i.e. (1.1) with potential (1.2)), and for (1.1) with potentials closely related to (1.2). Our results are summarized in Section 4.

2. The density and current correlations

The Fokker–Planck equation (1.1) relates to the overdamped Brownian motion dynamics of a classical gas with potential energy \(W\) in contact with a heat bath at inverse temperature \(\beta\). The function \(p\) is the p.d.f. for the event that the particles are at positions \(x_1, \ldots, x_N\) after time \(\tau\). As written in (1.1), the particles are assumed to be confined to a line, however the same equation applies in higher space dimensions if we simply replace \(x_j\) by \(x_j^{(s)}\) (the components of the position coordinate \(x_j\)) and assume summation over \(s\).
The dynamics specified by (1.1) is known [11,12] to be equivalent to the dynamics specified by the coupled Langevin equations

\[ \gamma \frac{dx_j(\tau)}{d\tau} = -\frac{\partial W}{\partial x_j} + \mathcal{F}_j(\tau) \quad (j = 1, \ldots, N) \] (2.1a)

where the random force \( \mathcal{F}_j(\tau) \) is a Gaussian random variable with zero mean and variance given by

\[ \overline{\mathcal{F}_j(\tau)\mathcal{F}_j(\tau')} = \frac{2\gamma}{\beta} \delta_{jj} \delta(\tau - \tau') \] (2.1b)

(the average denoted by an overline is with respect to the random force). In particular this means that the correlation functions can be calculated using either (1.1) or (2.1).

In the Fokker-Planck formalism the correlation functions can be specified in terms of the Green function \( G(x_1^{(0)}, \ldots, x_N^{(0)}|x_1, \ldots, x_N; \tau) \), which is by definition the solution of (1.1) subject to the initial condition

\[ p(x_1, \ldots, x_N; \tau = 0) = \prod_{j=1}^{N} \delta(x_j - x_j^{(0)}) \] (2.2)

Thus for observables \( A_\tau \) and \( B_\tau \) (e.g. \( A_\tau = n_\tau(x) := \sum_{j=1}^{N} \delta(x - x_j(\tau)) \), which corresponds to the microscopic density) and initial p.d.f. for the position of the particles \( f \), the correlation between \( A_{\tau_a} \) and \( B_{\tau_b} \) is defined as

\[ \langle A_{\tau_a}B_{\tau_b}\rangle^T = \langle A_{\tau_a}\rangle \langle B_{\tau_b}\rangle - \langle A_{\tau_a}B_{\tau_b}\rangle \] (2.3a)

where

\[ \langle A_{\tau_a}B_{\tau_b}\rangle = \int \int dx_1^{(0)} \cdots \int dx_N^{(0)} f(x_1^{(0)}, \ldots, x_N^{(0)}) \]
\[ \times \int \int dx_1^{(1)} \cdots \int dx_N^{(1)} A\{x_1^{(1)}\}G(x_1^{(0)}, \ldots, x_N^{(0)}|x_1^{(1)}, \ldots, x_N^{(1)}; \tau_a) \]
\[ \times \int \int dx_1^{(2)} \cdots \int dx_N^{(2)} B\{x_2^{(2)}\}G(x_1^{(1)}, \ldots, x_N^{(1)}|x_1^{(2)}, \ldots, x_N^{(2)}; \tau_b - \tau_a) \] (2.3b)

\[ \langle A_{\tau_a}\rangle = \int \int dx_1^{(0)} \cdots \int dx_N^{(0)} f(x_1^{(0)}, \ldots, x_N^{(0)}) \int dx_1 \cdots \int dx_N A\{x_j\} \]
\[ \times G(x_1^{(0)}, \ldots, x_N^{(0)}|x_1, \ldots, x_N; \tau_a), \] (2.3c)

and similarly the definition of \( \langle B_{\tau_b}\rangle \). The quantity \( \langle A_{\tau_a}B_{\tau_b}\rangle \) is referred to as the distribution function.
In the Langevin equation formalism the correlation between \( A_t \) at time \( \tau = 0 \) and \( B_t \) at time \( \tau_b \) is defined as

\[
\langle A_0 B_{\tau_b} \rangle_0 = \langle A_0 B_{\tau_b} \rangle_0 - \langle A_0 \rangle_0 \langle B_{\tau_b} \rangle_0 \tag{2.4}
\]

where \( \langle \rangle_0 \) denotes the average over the initial particle distribution, while the overline denotes the average with respect to the random force for times from 0 to \( \tau \). The equivalence between the Langevin and Fokker–Planck formalisms implies

\[
\langle A_0 B_{\tau_b} \rangle = \langle A_0 B_{\tau_b} \rangle_0 \tag{2.5}
\]

When using the Fokker–Planck formalism we will make use of the general fact [11] that

\[
e^{\beta W/2} \mathcal{L} e^{-\beta W/2} = \sum_{j=1}^{N} \left( \frac{1}{\beta} \frac{\partial^2}{\partial \lambda_j^2} - \frac{\beta}{4} \left( \frac{\partial W}{\partial \lambda_j} \right)^2 + \frac{1}{2} \frac{\partial^2 W}{\partial \lambda_j^2} \right)
\]

\[
= -\frac{1}{\beta} \sum_{j=1}^{N} \Pi_j^\dagger \Pi_j \tag{2.6a}
\]

where

\[
\Pi_j := \frac{1}{i} \frac{\partial}{\partial x_j} - \frac{\beta}{2} \frac{\partial W}{\partial x_j} . \tag{2.6b}
\]

Note that

\[
\Pi_j e^{-\beta W/2} = 0 . \tag{2.7}
\]

The formula (2.6) shows that after conjugation with \( e^{-\beta W/2} \) the Fokker–Planck operator transforms into a Hermitian operator. For some particular \( W \), which include (1.2), we have

\[
e^{\beta W/2} \mathcal{L} e^{-\beta W/2} = -\frac{1}{\beta} (H - E_0) \tag{2.8}
\]

where \( H \) is the Schrödinger operator for a quantum mechanical system with one and two body interactions only, and \( E_0 \) is the corresponding ground state energy. For example, with \( W \) given by (1.2a) Eq. (2.8) holds with

\[
H = -\sum_{j=1}^{N} \frac{\partial^2}{\partial x_j^2} + \frac{\beta^2}{4} \sum_{j=1}^{N} x_j^2 + \beta(\beta/2 - 1) \sum_{1 \leq j < k \leq N} \frac{1}{(x_j - x_k)^2} . \tag{2.9}
\]

Note from (2.6a) and (2.8) that the ground state is (up to normalization) given by

\[
\psi_0 = e^{-\beta W/2} . \tag{2.10}
\]
2.1. The current-current correlation

In the Langevin description of Brownian dynamics the current-current correlation is defined by choosing the observables $A_t$ and $B_t$ in (2.4) as the classical current:

$$A_0 = j_0(x_a), \quad B_t = j_t(x_b), \quad j_t(x) := \sum_{j=1}^{N} \frac{dx_j(t)}{dt} \delta(x - x_j(t)). \quad (2.11)$$

In the Fokker–Planck description the classical current has no immediate meaning as the velocities do not explicitly occur in (1.1). To define the current in this situation we make use of the formula (2.6a) and rewrite (2.3) in terms of time-dependent operators.

For this purpose we note from (2.2) that

$$G(x_1^{(0)}, \ldots, x_N^{(0)}; x_1^{(1)}, \ldots, x_N^{(1)}; \tau) = e^{\frac{\mathcal{W}}{2}} \prod_{i=1}^{N} \delta(x_i^{(1)} - x_i^{(0)}) \quad (2.12)$$

where the Fokker–Planck operator acts on $\{x_i^{(1)}\}$ only. Substituting (2.12) in (2.3b) allows the integration over $\{x_i^{(0)}\}$ to be carried out. Then substituting (2.12), with $\{x_i^{(0)}\}$, $\{x_i^{(1)}\}$ replaced by $\{x_i^{(1)}\}$, $\{x_i^{(2)}\}$, in the result and integrating over $\{x_i^{(1)}\}$ we obtain

$$\langle A_{\tau_a} B_{\tau_b} \rangle = \int dx_1^{(2)} \ldots \int dx_N^{(2)} B_{\tau_b} e^{\frac{\mathcal{W}}{2} e^{\frac{\mathcal{W}}{2}}} \int A_{\tau_a} e^{\frac{\mathcal{W}}{2} e^{\frac{\mathcal{W}}{2}}} f(x_1^{(2)}, \ldots, x_N^{(2)}). \quad (2.13)$$

Now it follows from (2.6a) that

$$e^{\frac{\mathcal{W}}{2}} = e^{\frac{-\mathcal{W}}{2}} e^{-\frac{1}{2} \sum_{i=1}^{N} \eta_{i} \bar{\eta}_{i} e^{\frac{\mathcal{W}}{2}}}, \quad (2.14)$$

so we can rewrite (2.13) as

$$\langle A_{\tau_a} B_{\tau_b} \rangle = \int dx_1^{(2)} \ldots \int dx_N^{(2)} e^{-\beta \frac{\mathcal{W}}{2}} B(\tau_b) A(\tau_a) f e^{\beta \frac{\mathcal{W}}{2}} \quad (2.15a)$$

where

$$A(\tau) := e^{\frac{1}{2} \sum_{i=1}^{N} \eta_{i} \bar{\eta}_{i}} e^{-\frac{1}{2} \sum_{i=1}^{N} \eta_{i} \bar{\eta}_{i}}$$

and similarly the definition of $B(\tau)$ (in deriving (2.15a) we have used the fact that $e^{\tau \sum_{i=1}^{N} \eta_{i} \bar{\eta}_{i}} e^{-\beta \frac{\mathcal{W}}{2}} = 1$, which follows from (2.7)).

Using (2.15b) we can define the current $j_{\tau}(x)$ by the continuity equation

$$\frac{\partial}{\partial \tau} n(x, \tau) = - \frac{\partial}{\partial x} j(x, \tau), \quad (2.16)$$

where $n(x, \tau)$ is defined by the r.h.s. of (2.15b) with

$$A_{\tau} = n_{\tau}(x) := \sum_{j=1}^{N} \delta(x - x_j(\tau)) \quad (2.17)$$
which is the microscopic density. Now (2.15b) is the imaginary time quantum mechanical formula for the evolution of the operator \(A_t\) in the Heisenberg picture. Further the Hamiltonian \(\sum_{j=1}^{N} \Pi_j^\dagger \Pi_j\) is of the form

\[
- \sum_{j=1}^{N} \frac{\partial^2}{\partial x_j^2} + V(x_1, \ldots, x_N),
\]

so we conclude from the usual quantum mechanical verification of the continuity equation that the sought definition of the current is

\[
j_i(x) = -\frac{i}{\gamma \beta} \sum_{j=1}^{N} \left( \frac{1}{i} \frac{\partial}{\partial x_j} \delta(x - x_j(\tau)) + \delta(x - x_j(\tau)) \frac{1}{i} \frac{\partial}{\partial x_j} \right).
\]

As written (2.19) assumes the particles are on a line. However the same formula applies for each component of the current vector in higher dimensions if we simply replace \(j_i(x)\) by \(j_i^{(x)}(x)\) and \(\partial/\partial x_j\) by \(\partial/\partial x_j^{(x)}\).

Note that with \(f\) proportional to \(e^{-\beta W}\) (the equilibrium state) in (2.15a) we can write

\[
\langle A_{x_a} B_{x_b} \rangle = \langle \psi_0 | B(\tau_b) A(\tau_a) | \psi_0 \rangle
\]

which is precisely the quantum mechanical formula (with \(\tau = it\beta\gamma\)) for ground state correlations. This shows that, up to a factor of \((i/\gamma \beta)^2\), the current–current distribution for Brownian motion described by the Fokker–Planck equation (2.1) with initial p.d.f. given by the equilibrium distribution is identical to the current–current distribution of the corresponding quantum mechanical system (recall (2.8)). The equivalence between the density–density distributions in this situation has previously been shown by Beenakker and Rejaei [1].

2.2. \(f\)-sum rule

Denote the density–density correlation by \(S\) so that

\[
S((x_a, 0), (x_b, \tau_b)) = \left\langle \sum_{j=1}^{N} \delta(x_a - x_j(0)) \sum_{k=1}^{N} \delta(x_b - x_k(\tau_b)) \right\rangle^T
\]

where the average on the r.h.s. is defined according to (2.3). When the initial distribution \(f\) in (2.2) is proportional to the equilibrium distribution \(e^{-\beta W}\) we have just remarked that \(S\) is identical (with \(\tau = it\beta\gamma\)) to the ground state density–density correlation of the quantum system with Hamiltonian \(\sum_{j=1}^{N} \Pi_j^\dagger \Pi_j\). Since this Hamiltonian is of the form (2.18) it follows that \(S\) must satisfy the so called \(f\)-sum rule [10]:

\[
\frac{\partial}{\partial \tau} \tilde{S}(k, \tau) \bigg|_{\tau=0} = -\frac{k^2}{\gamma \beta} P_k
\]

(2.22a)
where

\[ \tilde{S}(k, \tau) := \int_{-\infty}^{\infty} S((x_a, 0), (x_b, \tau)) e^{ik(x_a-x_b)} d(x_a-x_b) \] (2.22b)

and it is assumed that the ground state is translationally invariant so that \( S \) depends on \( x_a - x_b \). Subject to this assumption (2.22a) is valid for \( N \) finite as well as in the thermodynamic limit.

In fact it is possible to derive and indeed generalize the \( f \)-sum rule entirely within the Brownian motion setting using either the Fokker–Planck or Langevin equation formalism (for definiteness we will consider the latter). The generalization is that we can take the initial distribution as proportional to the Boltzmann factor \( e^{-\beta'W} \) (where we may have \( \beta' \neq \beta \)) and we do not assume translational invariance.

We consider the quantity

\[ \tilde{S}(x, k, \tau) := \int_{-\infty}^{\infty} S((x, 0), (x_b, \tau)) e^{ikx_b} dX_b, \] (2.23)

where \( S \) is given by (2.20) with the average now computed according to the r.h.s. of (2.5) because we will use the Langevin formalism. Differentiating with respect to \( \tau \) gives

\[ \gamma \frac{\partial \tilde{S}(x, k, \tau)}{\partial \tau} = ik \left( \sum_{j=1}^{N} \delta(x - x_j(0)) \sum_{l=1}^{N} \gamma \dot{x}_l(\tau) e^{ikx_l(\tau)} \right). \] (2.24)

As \( \tau \) approaches zero, we see from the Langevin equation (2.1) that

\[ \gamma \dot{x}_j(\tau) = -\frac{\partial W}{\partial x_j} \bigg|_{\tau=0} + \mathcal{F}_j(\tau) + O(\tau) \] (2.25a)

and so

\[ e^{ikx_j(\tau)} = e^{ikx_j(0)} \left( 1 + \frac{ik}{\gamma} \int_{0}^{\tau} \mathcal{F}_j(\tau') d\tau' + O(\tau) \right) \] (2.25b)

where the term proportional to \( \tau \) not explicitly written is independent of \( \mathcal{F}_j(\tau) \). Substituting (2.25) in (2.24) and using (2.1b) to compute the average with respect to the random force gives

\[ \gamma \frac{\partial \tilde{S}(x, k, \tau)}{\partial \tau} \bigg|_{\tau=0} = ik \sum_{j,l=1}^{N} \left( \delta(x - x_j(0)) e^{ikx_j(\tau)} \left( -\frac{\partial W}{\partial x_j} \bigg|_{\tau=0} + \frac{2ik}{\beta} \int_{0}^{\tau} \delta(\tau - \tau') d\tau \right) \right). \] (2.26)
The delta function in (2.26) is non-zero on the boundary of the interval of integration so we take one half of the value of the integrand. After averaging over the initial distribution, the corresponding term of (2.26) is then identified as proportional to \( \tilde{S}(x,k,0) \). To simplify the term involving \( \partial W/\partial x_j \) in (2.26) we recall that the initial distribution is assumed proportional to \( e^{-\beta W} \). Since

\[
-\frac{\partial W}{\partial x_j} e^{-\beta W} = \frac{1}{\beta'} \frac{\partial}{\partial x_j} e^{-\beta' W}
\]

integration by parts gives that this term can be rewritten as

\[
-\frac{ik}{\beta'} \tilde{S}(x,k,0) + \frac{1}{\beta'} \frac{\partial}{\partial x} (e^{ikx} \rho(x)) .
\]  (2.27)

Combining these results gives for our generalization of the \( f \)-sum rule

\[
\frac{\partial \tilde{S}(x,k,\tau)}{\partial \tau} \bigg|_{\tau=0} = -k^2 \left( \frac{1}{\beta' \gamma} - \frac{1}{\beta' \gamma} \right) \tilde{S}(x,k,0) + \frac{ik}{\beta' \gamma} \frac{\partial}{\partial x} (e^{ikx} \rho(x)) ,
\]  (2.28)

which reduces to (2.22a) in the case \( \beta' = \beta, \rho(x) = \rho \).

We emphasize that (2.28) holds for a finite system, as well as in the thermodynamic limit. It has been presented in the one-dimensional case, but by introducing the higher-dimensional Fourier transform in (2.23) and repeating the working we see that (2.28) holds with \( k^2 \) replaced by \( \sum_{x} (k^{(x)})^2 \) in the first term and \( k, \partial/\partial x, kx \) replaced by \( k^{(x)}, \partial/\partial x^{(x)}, k^{(x)} x^{(x)} \) (with summation over \( x \)) in the last term.

### 2.3. Static current–current distribution

For the particular Schrödinger operator

\[
H = -\sum_{j=1}^{N} \frac{\partial^2}{\partial x_j^2} + \beta(\beta/2 - 1) \left( \frac{\pi}{L} \right)^2 \sum_{1 \leq j < k \leq N} \frac{1}{\sin^2 \pi(x_j - x_k)/L} , \quad 0 \leq x_j \leq L ,
\]  (2.29)

which is related via Eq. (2.8) to the Fokker–Planck operator (1.1) with potential (1.2b), it has been shown by Taniguchi et al. [3] that the static current–current distribution is given in terms of the static density–density distribution by

\[
\langle \psi_0 | j(x_b,0) j(x_a,0) | \psi_0 \rangle = -\frac{\beta(\pi/L)^2}{\sin^2 \pi(x_b - x_a)/L} \langle \psi_0 | \rho(x_b,0) \rho(x_a,0) | \psi_0 \rangle , \quad x_a \neq x_b .
\]  (2.30)

From the remarks below (2.20), (2.30) is equivalent to the statement that for the Fokker–Planck system with potential (1.2b) and initial p.d.f. equal to the equilibrium
distribution,

\[ \langle j_0(x_a)j_0(x_b) \rangle = -\frac{\beta(\pi/L)^2}{\sin^2 \pi(x_b - x_a)/L} \left( \frac{i}{\gamma\beta} \right)^2 \langle n_0(x_b)n_0(x_a) \rangle. \]  

(2.31)

Eq. (2.30) was derived in Ref. [3] using the factorization (2.6a). The same method readily extends to provide an analogous result for all Fokker–Planck systems in which \( W \) consists of one and two body potentials only.

Thus we are considering the average (2.15a) with \( \tau_a \) and \( \tau_b \) equal to zero, the observables \( A \) and \( B \) given by the current (2.19) and the initial distribution \( f \) proportional to \( e^{-\beta W} \). To apply the method of Ref. [3], we substitute for \( \partial / \partial x_j \) in (2.19) using \( \eta_j \) as defined by (2.6b) to obtain

\[ j_0(x) = -\frac{i}{\beta\gamma} \sum_{j=1}^{N} (\Pi_j^\dagger \delta(x - x_j) + \delta(x - x_j)\Pi_j). \]  

(2.32)

From this formula and (2.7) we have

\[
\langle j_0(x_a)j_0(x_b) \rangle = \left( \frac{i}{\beta\gamma} \right)^2 \sum_{j,k=1}^{N} \frac{1}{Z} \int dx_1^{(2)} \cdots \int dx_N^{(2)} e^{-\beta W/2} \delta(x_b - x_j)\Pi_j^\dagger \delta(x_a - x_k)e^{-\beta W/2}
\]

(2.33)

where

\[ Z := \int dx_1^{(2)} \cdots \int dx_N^{(2)} e^{-\beta W}. \]

Let us now suppose \( x_a \neq x_b \). Then the \( j = k \) term vanishes. But for \( j \neq k \) we have

\[ [\Pi_j, \Pi_k^\dagger] = \beta \frac{\partial^2 W}{\partial x_j \partial x_k}. \]  

(2.34)

Using this formula, the fact that \( \Pi_j \delta(x - x_k) = \delta(x - x_k)\Pi_j \) for \( j \neq k \), the property (2.7) and the definition (2.17) we deduce from (2.33) that

\[ \langle j_0(x_a)j_0(x_b) \rangle = \left( \frac{i}{\beta\gamma} \right)^2 \beta \frac{\partial^2 W}{\partial x_1 \partial x_2} \bigg|_{x_1 = x_a, x_2 = x_b} \langle n_0(x_a)n_0(x_b) \rangle, \]  

(2.35)

which is the sought formula. We remark that it is also possible to derive (2.35) within the Langevin equation setting used in the derivation of the \( f \)-sum rule given in the previous section. Also, for \( W \) given by (1.2b), note that (2.35) reduces to (2.31).

Our derivation has been presented in the one-dimensional case. The same result applies in higher dimensions for the correlations between components \( j_0^{(\sigma)}(x_a), j_0^{(\beta)}(x_b) \) of the current vector provided the \( x_1, x_a \) and \( x_2, x_b \) outside the average on the r.h.s. are the components \( (\sigma) \) and \( (\beta) \), respectively, of the corresponding position vector.
2.4. Small-\(k\) behaviour of structure factor

In this section we will consider Brownian motion described by the Fokker–Planck equation (1.1), with the potential \(W\) consisting of general one and two body terms

\[
W = \sum_{j=1}^{N} V_1(x_j) + \sum_{1 \leq j < k \leq N} V_2(|x_j - x_k|) \tag{2.36}
\]

such that the potential \(V_2\) is non-integrable at infinity. For the particular case of this type

\[
V_2(x) = -\log |x|. \tag{2.37}
\]

Dyson [5] introduced the hydrodynamic equation

\[
\gamma j(x, \tau) = -\rho(x, \tau) \frac{\partial}{\partial x} \left( V_1(x) - \int_{-\infty}^{\infty} dx' \rho(x', \tau) V_2(|x - x'|) \right) \tag{2.38}
\]

(note that the r.h.s. represents the force density) to study the large-wavelength density fluctuations of the system. Here \(j(x, \tau)\) and \(\rho(x, \tau)\) represent smoothed, continuum approximations to the microscopic current and density.

In the case (2.37) it has been deduced in Ref. [1] that for a uniform initial state

\[
\rho(x', 0) = \rho \tag{2.39}
\]

the Fourier transform of the density–density correlation defined by (2.22) has the small-\(k\) behaviour

\[
\tilde{S}(k, \tau) \sim \tilde{S}(k, 0) e^{-k^2 \rho \tilde{V}_2(k)/\gamma} \tag{2.40}
\]

where

\[
\tilde{V}_2(k) := \int_{-\infty}^{\infty} V_2(|x|) e^{ikx} \, dx. \tag{2.41}
\]

Inspection of the derivation given in [1] shows (2.40) to be a consequence of (2.38) independent of the particular \(V_2\), so its validity for general \(V_2\) is tied to the asymptotic validity of the hydrodynamic approximation (2.38). We expect (2.38) to correctly describe the large-wavelength density fluctuations whenever the potential \(V_2(|x|)\) is not integrable at infinity (for integrable potentials the ‘electric field term’ involving the integral on the r.h.s. of (2.38) is not present, rather the force density is due to a pressure gradient; see the next subsection). Now for potentials not integrable at infinity it is expected [13] that

\[
\tilde{S}(k, 0) \sim 1/\beta' \tilde{V}_2(k) \tag{2.42}
\]
(here \( \beta' \) is used as in the discussion of the generalized \( f \)-sum rule, and thus may be different to \( \beta \) which occurs in (1.1)) and so (2.40) reads
\[
\tilde{S}(k, \tau) \sim \frac{1}{\beta' \tilde{V}_2(k)} e^{-k^2 \tau \rho \tilde{V}_2(k) / \gamma}.
\] (2.43)

Note that this expression is consistent with the generalized \( f \)-sum rule (2.28) with \( \rho(x) = \rho \) (note that the first term in (2.28) is of a lower order than the second since, as is seen from (2.42), \( \tilde{S}(k, 0) \to 0 \) as \( k \to 0 \) whenever \( V_2(|x|) \) is not integrable at infinity). It is expected to be valid for non-integrable potentials in higher dimensions provided (2.41) is suitably modified.

One consequence of (2.43) is the sum rule
\[
\int_0^\infty \tilde{S}(k, \tau) d\tau \sim \frac{\gamma}{\rho \beta' k^2 (\tilde{V}_2(k))^2} \quad \text{as } k \to 0.
\] (2.44)

In the particular case when \( V_2 \) is given by (2.37) and \( \beta' = \beta \), we noted in Section 2.1 that \( \tilde{S}(k, \tau) \) in the Fokker–Planck system is identical to \( S(k, \tau) \) for the ground state of the Calogero–Sutherland quantum system with Schrödinger operator (2.9). Now for a quantum fluid the integral (2.44) is related to the compressibility of the ground state. This is consistent with the fact that for \( V_2 \) given by (2.37)
\[
\tilde{V}_2(k) \sim \pi / |k|
\] (2.45)

and so the r.h.s. of (2.44) is a constant.

Another consequence of (2.43) is the evaluation of the large-\( \tau \) mean square displacement of a particle from its initial position. Thus from Ref. [14], assuming this displacement diverges, \( \tilde{S}(k, 0) \) diverges at the origin and that the state is homogeneous, we have for large-\( \tau \)
\[
\langle (x(\tau) - x(0))^2 \rangle \sim \frac{2}{\pi \rho^2} \int_0^{\rho c} \frac{dk}{k^2} (\tilde{S}(k, 0) - \tilde{S}(k, \tau))
\] (2.46)

where \( c \ll 1 \) is some constant. To make further progress, suppose \( V_2(r) \sim r^{-\alpha}, \; 0 < \alpha < 1 \) as \( r \to \infty \), so that [15]
\[
\tilde{V}_2(k) \sim \frac{\pi |k|^\alpha}{\Gamma(\alpha) \cos \pi \alpha / 2} \quad \text{as } k \to 0.
\] (2.47)

The bounds on \( \alpha \) are required for the validity of (2.46). Indeed the bound \( \alpha < 1 \) is required so that \( \tilde{S}(k, 0) \) diverges at the origin, while substituting (2.47) in (2.43) and then substituting the result in (2.46) and shows that for (2.46) to diverge as \( \tau \to \infty \) we must have \( \alpha > 0 \). To calculate the large-\( \tau \) behaviour we make the substitutions of the previous sentence and change variables \( y = k^{\alpha+1} (\pi / \Gamma(\alpha) \cos \pi \alpha / 2) \tau \rho / \gamma \). This gives
\[
\langle (x(\tau) - x(0))^2 \rangle \sim c(\beta', \gamma, \rho)^{\alpha/(\alpha+1)} \quad \text{as } \tau \to \infty
\] (2.48a)
where
\[
\begin{align*}
c(\beta', \gamma, \rho) &= \frac{2}{\beta' \pi \rho^2} \left( \frac{\rho \Gamma(\gamma)}{\pi} \right)^{\frac{\gamma}{\gamma+1}} \left( \frac{\Gamma(\gamma) \cos \pi \gamma / 2}{\pi} \right)^{1/(\gamma+1)} \\
&\times \int_0^\infty y^{-2+1/(1+\gamma)}(1-e^{-y})dy, \quad (2.48b)
\end{align*}
\]

which is the sought formula. In the case \(\gamma > 1\) (integrable potential) it is shown in Ref. [14] that \(\langle (x(\tau) - x(0))^2 \rangle\) is proportional to \(\tau^{1/2}\), independent of \(\alpha\).

For a one-dimensional translationally invariant initial state which is the equilibrium state, the small-\(k\) behaviour (2.43) also implies the small-\(k\) behaviour of the Fourier transformed current-current correlation
\[
\bar{C}(k, \tau) := \int_{-\infty}^\infty \langle j_0(0) j_\tau(x) \rangle e^{ikx} dx.
\]

To see this, we recall [1] that for a one-dimensional translationally invariant system, a consequence of the continuity equation (2.16) is that
\[
\frac{\partial}{\partial \tau_a} \hat{S}(k; \tau_a, \tau_b) \bigg|_{\tau_a = 0, \tau_b = \tau} = \frac{\partial^2}{\partial \tau_a \partial \tau_b} \hat{S}(k; \tau_a, \tau_b), \quad (2.49)
\]

where
\[
\hat{S}(k; \tau_a, \tau_b) := \int_{-\infty}^\infty S(0, \tau_a, (x, \tau_b)) e^{ikx} dx.
\]

But by the assumption that the initial state is the equilibrium state \(\hat{S}(k; \tau_a, \tau_b) = \hat{S}(k, \tau_b - \tau_a)\) and so (2.49) reads
\[
\bar{C}(k, \tau) = -\frac{1}{k^2} \frac{d^2}{d\tau^2} \hat{S}(k, \tau), \quad (2.50)
\]

Substituting (2.50) in (2.43) (with \(\beta' = \beta\)) gives for the small-\(k\) asymptotic formula
\[
\bar{C}(k, \tau) \sim -\frac{\rho^2}{\beta \gamma^2} k^2 \tilde{V}_2(k) e^{-k^2 \tau \rho^2(k)/\gamma}. \quad (2.51)
\]

For the Dyson log-gas, when \(\tilde{V}_2(k)\) is given by (2.45), (2.51) gives the \(\tau = 0\) result
\[
\bar{C}(k, 0) \sim -\rho^2 \pi |k|/\beta \gamma^2.
\]

By taking the inverse transform we deduce the large-\(x_{ab}\) behaviour
\[
\langle j_0(x_a) j_0(x_b) \rangle \sim \frac{\rho^2}{\beta \gamma^2 (x_a - x_b)^2}, \quad (2.52)
\]
which is consistent with the large-\(x_{ab}\) behaviour of the thermodynamic form of (2.31), since in this limit \(\langle n_0(x_b)n_0(x_a) \rangle \sim \rho^2\).

We also have the analogue of (2.44) in the setting of the validity of (2.50). Thus integrating (2.50) with respect to \(\tau\) gives

\[
\int_0^\infty \hat{C}(k, \tau) d\tau = -\frac{1}{k^2} \frac{d}{d\tau} \tilde{S}(k, \tau) \bigg|_{\tau=0} = \frac{\rho}{\gamma \beta},
\]

(2.53)

where to obtain the last equality the \(f\)-sum rule (2.22a) has been used. Note that unlike (2.44), Eq. (2.53) is valid for all values of \(k\).

2.5. Second-order correction

To obtain the next order correction in the small-\(k\) expansion (2.40) we refine the hydrodynamic equation (2.38) to include a pressure gradient \(-\partial p(x, \tau)/\partial x\) as an additional force density on the r.h.s. The strategy for the solution of this equation is the same as given in [1] for the solution of (2.40). The first step is to differentiate both sides of the equation with respect \(x\) and to then substitute for \(\partial j(x, \tau)/\partial x\) using the continuity equation (2.16). After linearizing in \(\delta p(x, \tau) := p(x, \tau) - \rho\) (note that the pressure is linearized by \(p(x, \tau) - \rho = \frac{\partial p}{\partial \rho} \delta \rho(x, \tau)\) where \(p\) denotes the bulk pressure) and assuming the equilibrium condition that for large-\(L\)

\[
\frac{\partial}{\partial x} \left( V_1(x) - \rho \int_{-L}^L V_2(|x - x'|) dx' \right) = 0
\]

we obtain

\[
\frac{\partial \delta p(x, \tau)}{\partial \tau} = \frac{\rho}{\gamma} \frac{\partial^2}{\partial x^2} \int_{-\infty}^\infty V_2(|x - x'|) \delta \rho(x', \tau) dx' + \frac{1}{\gamma} \frac{\partial p}{\partial \rho} \frac{\partial^2}{\partial x^2} \delta \rho(x, \tau).
\]

Taking the Fourier transform of both sides and integrating by parts twice gives a simple differential equation for \(\delta \tilde{\rho}(k, \tau)\) which has solution

\[
\delta \tilde{\rho}(k, \tau) = \delta \tilde{\rho}(k, 0) \exp \left( - \left( k^2 \rho \tilde{V}_2(k) + \frac{\partial p}{\partial \rho} k^2 \right) \frac{\tau}{\gamma} \right).
\]

(2.54)

But

\[
\tilde{S}(k, \tau) = \frac{1}{L} \langle \delta \tilde{\rho}(-k, 0) \delta \tilde{\rho}(k, \tau) \rangle.
\]

(2.55)

Substituting (2.54) gives the desired second-order correction to (2.40):

\[
\tilde{S}(k, \tau) \sim \tilde{S}(k, 0) \exp \left( - \left( k^2 \rho \tilde{V}_2(k) + \frac{\partial p}{\partial \rho} k^2 \right) \frac{\tau}{\gamma} \right).
\]

(2.56)
Note that for consistency with the second term of the \( f \)-sum rule (2.28) with \( \rho(x) = \rho \), which gives the leading small-\( k \) behaviour, we must have

\[
\tilde{S}(k, 0) \sim \frac{1}{\beta'(\tilde{V}_2(k) + \frac{1}{\rho} \frac{\partial \rho}{\partial \rho})}.
\]

(2.57)
A thermodynamic derivation of this result is given in the Appendix.

3. Exact results

3.1. GOE \( \rightarrow \) GUE transition

Consider the Dyson Brownian motion model specified by the Fokker-Planck equation (1.1) with \( \beta = 2 \), \( W \) given by (1.2a) and the initial distribution proportional to \( e^{-W} \). Since \( e^{-W} \) is proportional to the eigenvalue p.d.f. for the Gaussian Orthogonal Ensemble (GOE) of random real symmetric matrices, while \( e^{-2W} \) is proportional to the eigenvalue p.d.f. for the Gaussian Unitary Ensemble (GUE) of random Hermitian matrices, this specific Dyson Brownian motion model describes the GOE \( \rightarrow \) GUE transition. In the \( N \to \infty \) limit (suitably scaled) the system is translationally invariant with bulk density \( \rho \) and the Fourier transform of the density–density correlation (2.22) is given exactly by [4, Eq. (3.32) with the correction that an additional factor \( e^{(\tau \pi |k|/\rho/\gamma)(1-|k|/2\pi \rho)} \) be included in the first term of (3.32c)]

\[
\tilde{S}(k, \tau) = \frac{2e^{\gamma}}{\pi \tau |k|} \exp(-\tau \pi |k|/\rho/\gamma) \sinh(\tau k^2/2\gamma) - \left| k \right| e^{\tau k^2/2\gamma} \int_1^{1+|k|/\pi \rho} dk_1 \frac{1}{k_1} e^{-\pi \rho |k_1|/\gamma}, \quad 0 \leq |k| \leq 2\pi \rho \quad (3.1a)
\]

\[
\tilde{S}(k, \tau) = \frac{2e^{\gamma}}{\pi \tau |k|} \exp(-\tau k^2/2\gamma) \sinh(\tau \pi |k|/\rho/\gamma) - \left| k \right| e^{\tau k^2/2\gamma} \int_{-1+|k|/\pi \rho}^{1+|k|/\pi \rho} dk_1 \frac{1}{k_1} e^{-\pi \rho |k_1|/\gamma}, \quad |k| \geq 2\pi \rho \quad (3.1b)
\]

This result provides an illustration of the generalized \( f \)-sum rule (2.20) in the case \( \beta \neq \beta' \). Thus according to (2.28) with \( \rho(x) = \rho \), \( \beta = 2 \), \( \beta' = 1 \) and \( x = 0 \), for all \( k \) (3.1) must satisfy

\[
\frac{\partial \tilde{S}(k, \tau)}{\partial \tau}\bigg|_{\tau=0} = \frac{k^2}{2\gamma} \tilde{S}(k, 0) - \frac{k^2 \rho}{\gamma}. \quad (3.2a)
\]

An elementary calculation verifies that this is indeed so. Also, the small-\( k \) expansion of (3.1a) is consistent with (2.56) and (2.57) with \( \beta' = 1 \), \( \tilde{V}_2(k) \) given by (2.45) and \( \frac{\partial \rho}{\partial \rho} \) calculated from the equation of state \( \beta' \rho = (1 - \beta'/2) \rho \).
3.2. Free fermion type system

As remarked earlier, for a number of special choices of $W$ including (1.2), Eq. (2.8) holds relating the Fokker–Planck operator to a Schrödinger operator. A further property of this Schrödinger operator, as is illustrated by (2.9), is that at the special coupling $\beta = 2$ the coefficient of the two-body term vanishes and the system can be regarded as free fermions in an external potential.

Consider such a situation in which the one-particle Schrödinger operator, $H_1$ say, has a complete set of real, orthonormal eigenfunctions \( \{\psi_k\}_{k=0,1,...} \) and corresponding eigenvalues \( \{\varepsilon_k + E_0\}_{k=0,1,...} \) and the particles are confined to an interval $I$. A standard calculation (see e.g. Ref. [1]) gives that the density–density correlation for the $N$-particle system is given by

\[
S((x,0),(x',\tau)) = \sum_{p=0}^{\infty} \psi_p(x) \psi_p(x') e^{-\varepsilon_p \tau / \gamma \beta} \sum_{q=0}^{N-1} \psi_q(x) \psi_q(x') e^{\varepsilon_q \tau / \gamma \beta} 
- \sum_{p=0}^{N-1} \psi_p(x) \psi_p(x') e^{-\varepsilon_p \tau / \gamma \beta} \sum_{q=0}^{N-1} \psi_q(x) \psi_q(x') e^{\varepsilon_q \tau / \gamma \beta} .
\]  

(3.2b)

Since here we have set $t = \tau / (\beta \gamma)$, this represents the density–density correlation for the corresponding Fokker–Planck system with initial state equal to the equilibrium state. This exact result can be used to illustrate the generalized $f$-sum rule (2.20) in the case of finite $N$ and non-uniform density $\rho(x)$.

Thus we want to show that (3.2) satisfies

\[
\frac{\partial S(x,k,\tau)}{\partial \tau} \bigg|_{\tau=0} = \frac{ik}{\beta \gamma} \frac{\partial}{\partial x} (e^{ikx} \rho(x)) ,
\]  

(3.3)

where $\tilde{S}(x,k,\tau)$ is defined by (2.23) with $S$ therein given by (3.2), and $\rho(x)$ is given by

\[
\rho(x) = \sum_{p=0}^{N-1} (\psi_p(x))^2
\]  

(3.4)

(this latter formula follows from the free fermion calculation). For this purpose we note that differentiating the second term in (3.2) with respect to $\tau$ and then setting $\tau = 0$ gives zero as the two terms which result cancel. Using this feature and the fact $(H_1 - E_0) \psi_q(x') = \varepsilon_q \psi_q(x')$ we see that

\[
\frac{\partial}{\partial \tau} S((x,0),(x',\tau)) \bigg|_{\tau=0} = -\frac{1}{\gamma \beta} \sum_{p=0}^{\infty} \psi_p(x)((H_1 - E_0) \psi_p(x')) \sum_{q=0}^{N-1} \psi_q(x) \psi_q(x') 
+ \frac{1}{\gamma \beta} \sum_{p=0}^{\infty} \psi_p(x) \psi_p(x') \sum_{q=0}^{N-1} \psi_q(x)(H_1 - E_0) \psi_q(x') .
\]  

(3.5)
Taking the Fourier transform of both sides with respect to $x'$ and integrating by parts (twice) the first term on the r.h.s. using the explicit formula

$$H_1 = -\frac{d^2}{dx'^2} + V(x')$$

we see that the last term cancels. Furthermore, noting by completeness that

$$\sum_{\rho=0}^{\infty} \psi_{\rho}(x')\psi_\rho(x) = \delta(x' - x)$$

we obtain

$$\frac{\partial}{\partial \tau} S((x,0),(x',\tau)) \bigg|_{\tau=0} = \frac{2ik}{\gamma\beta} \int e^{ikx'} \delta(x - x') \frac{\partial}{\partial x'} \sum_{q=0}^{N-1} \psi_{\rho}(x)\psi_{\rho}(x') \, dx'$$

$$= \frac{k^2}{\gamma\beta} \int e^{ikx'} \delta(x - x') \sum_{q=0}^{N-1} \psi_{\rho}(x)\psi_{\rho}(x') \, dx'$$

$$= \frac{ik}{\gamma\beta} e^{ikx} \frac{d}{dx} \rho(x) - \frac{k^2}{\gamma\beta} e^{ikx} \rho(x)$$

(3.6)

where to obtain the last equality (3.4) has been used. Comparing (3.6) with the $f$-sum rule formula (3.3) shows that the required agreement with that result has been demonstrated.

3.3. The Dyson Brownian motion model for rational $\beta$

In the thermodynamic limit the density-density correlation $S(x,\tau)$ for the Dyson Brownian motion model specified by the Fokker–Planck equation (1.1) with $W$ given by (1.2b) has been calculated for all rational values of $\beta$. With $\beta/2 := \lambda = p/q$ ($p$ and $q$ relatively prime) the result is [9]

$$\lim_{N,L \to \infty} \frac{S(x,\tau)}{N/L} = C_{p,q}(\lambda) \prod_{i=1}^{q} \int_{0}^{\infty} dx_i \prod_{j=1}^{p} \int_{0}^{1} dy_j Q_{p,q}^2 F(q, p, \lambda | \{x_i, y_j\})$$

$$\times \cos Q_{p,q} x \exp(-E_{p,q} \tau/2\gamma)$$

(3.7a)

where the momentum $Q$ and the energy $E$ variables are given by

$$Q_{p,q} := 2\pi \rho \left( \sum_{i=1}^{q} x_i + \sum_{j=1}^{p} y_j \right)$$

$$E_{p,q} := (2\pi \rho)^2 \left( \sum_{i=1}^{q} \epsilon_p(x_i) + \sum_{j=1}^{p} \epsilon_H(y_j) \right)$$

(3.7b)
with
\[ \varepsilon_p(x) = x(x + \lambda) \quad \text{and} \quad \varepsilon_H(y) = \lambda y(1 - y), \] (3.7c)

the form factor \( F \) is given by
\[ F(q, p, \lambda; \{x_i, y_j\}) = \prod_{i=1}^{q} \prod_{j=1}^{p} (x_i + \lambda y_j)^{-2} \prod_{i<j} \frac{|x_i - x_j|^{2\lambda} \prod_{j<i} |y_j - y_{j'}|^{2/\lambda}}{\prod_{i=1}^{q} (\varepsilon_p(x_i))^{1-\lambda} \prod_{j=1}^{p} (\varepsilon_H(y_j))^{1-1/\lambda}} \] (3.7d)

and the normalization is given by
\[ C_{p,q}(\lambda) = \frac{\lambda^2 p^q (q-1)!^{1/2}}{2 \pi^2 p^q 1^{q-1}} \frac{\Gamma^q(\lambda) \Gamma^p(1/\lambda)}{\prod_{i=1}^{q} \Gamma^2(p - \lambda(i-1)) \prod_{j=1}^{p} \Gamma^2(1 - (j-1)/\lambda)}. \] (3.7e)

This exact result can be used to illustrate the sum rules (2.43) and (2.57). To begin, note that \( S(x, \tau) \) is even in \( x \) and so \( \tilde{S}(k, \tau) \) is even in \( k \). It therefore suffices to consider the case \( k > 0 \). With this assumption, from (3.7)
\[ \tilde{S}(k, \tau) = \pi C_{p,q}(\lambda) \prod_{i=1}^{q} \int_{0}^{\infty} dx_i \prod_{j=1}^{p} \int_{0}^{1} dy_j Q^2_{p,q} F(q, p, \lambda; \{x_i, y_j\}) \]
\[ \times \delta(k - Q_{p,q}) \exp(-Q_{p,q} \tau/2\lambda). \]

Changing variables \( x_i \rightarrow kx_i, \ y_j \rightarrow ky_j \) gives
\[ \tilde{S}(k, \tau) = \pi k C_{p,q}(\lambda) e^{-k\rho \tau/\gamma} \prod_{i=1}^{q} \int_{0}^{\infty} dx_i \prod_{j=1}^{p} \int_{0}^{1} dy_j Q^2_{p,q} \tilde{F}(q, p, \lambda; \{x_i, y_j\}; k) \]
\[ \times \delta(1 - Q_{p,q}) \exp(-\tilde{E}_{p,q,k} \tau/2\lambda), \] (3.8a)

where
\[ \tilde{F}(q, p, \lambda; \{x_i, y_j\}; k) = \prod_{i=1}^{q} \prod_{j=1}^{p} (x_i + \lambda y_j)^{-2} \frac{\prod_{i<j} |x_i - x_j|^{2\lambda} \prod_{j<i} |y_j - y_{j'}|^{2/\lambda}}{\prod_{i=1}^{q} (\lambda(kx_i + \lambda))^{1-\lambda} \prod_{j=1}^{p} (\lambda y_j(1 - y_{j'}))^{1-1/\lambda}} \] (3.8b)

and
\[ \tilde{E}_{p,q,k} = (2\pi \rho)^2 k^2 \left( \sum_{i=1}^{q} x_i^2 - \lambda \sum_{j=1}^{p} y_j^2 \right). \] (3.8c)

In the limit \( k \rightarrow 0 \) the integral in (3.8a) is independent of \( k \), so to leading order we have
\[ \tilde{S}(k, \tau) \sim \pi |k| C_{p,q}(\lambda) I(\lambda) e^{-|k| \rho \tau/\gamma} \] (3.9)
where

\[ I(\lambda) = \prod_{i=1}^{q} \int_{0}^{\infty} dx_i \prod_{j=1}^{p} \int_{0}^{\infty} dy_j Q_{p,q}^2 G(q, p, \lambda|\{x_i, y_j\}) \delta(1 - Q_{p,q}) \]  

(3.10)

with

\[ G(q, p, \lambda|\{x_i, y_j\}) = \prod_{i=1}^{q} \prod_{j=1}^{p} (x_i + \lambda y_j)^{-2} \frac{\prod_{i<j} |x_i - x_j|^2 \prod_{j<j'} |y_j - y_{j'}|^2}{\prod_{i=1}^{q} x_i^{1-\lambda} \prod_{j=1}^{p} \lambda y_j^{1-1/\lambda}}. \]  

(3.11)

To evaluate (3.10) we rewrite it as

\[ I(\lambda) = \lim_{\varepsilon \to 0^+} \frac{1}{2\pi} \int_{-\infty}^{\infty} du e^{iu} \prod_{i=1}^{q} \int_{0}^{\infty} dx_i \prod_{j=1}^{p} \int_{0}^{\infty} dy_j Q_{p,q}^2 G(q, p, \lambda|\{x_i, y_j\}) e^{-\varepsilon Q_{p,q}} e^{-iQ_{p,q}}. \]  

(3.12)

Now change variables

\[ x_i \rightarrow \frac{1}{\varepsilon + iu} x_i, \quad y_j \rightarrow \frac{1}{\varepsilon + iu} y_j. \]

This shows

\[ I(\lambda) = \left( \lim_{\varepsilon \to 0^+} \frac{1}{2\pi} \int_{-\infty}^{\infty} du \frac{e^{iu}}{(\varepsilon + iu)^2} \right) \prod_{i=1}^{q} \int_{0}^{\infty} dx_i \prod_{j=1}^{p} \int_{0}^{\infty} dy_j Q_{p,q}^2 G(q, p, \lambda|\{x_i, y_j\}) e^{-\varepsilon Q_{p,q}}. \]  

(3.13)

The first integral above equals unity in the limit \( \varepsilon \rightarrow 0^+ \), while the multiple integral has been evaluated in an earlier work [16] as equal to \( 1/(C_{p,q}(\lambda)\pi^2 \beta) \). Substituting in (3.13) and then substituting the result in (3.9) gives the hydrodynamic result (2.43) with the substitution (2.45).

It is also possible to illustrate (2.57) by expanding (3.8a) to the next order in \( k \):

\[ \tilde{F}(q, p, \lambda|\{x_i, y_j\}; k) \sim G(q, p, \lambda|\{x_i, y_j\})(1 + k(1 - 1/\lambda)Q_{p,q}/2\pi p). \]

This gives

\[ \tilde{S}(k, 0) \sim \frac{|k|}{n\beta} + \frac{1}{2\pi p} (1 - 1/\lambda)\pi^2 C_{p,q}(\lambda) \prod_{i=1}^{q} \int_{0}^{\infty} dx_i \prod_{j=1}^{p} \int_{0}^{\infty} dy_j Q_{p,q}^3 \]

\[ \times G(q, p, \lambda|\{x_i, y_j\}) \delta(1 - Q_{p,q}). \]  

(3.14)

Comparing the multiple integral above with that in (3.10) shows that they are identical apart the factor of \( Q_{p,q}^3 \) which reads \( Q_{p,q}^2 \) in (3.10). But this factor does not effect the
value of the integral due to factor of $\delta(1 - Q_{p,q})$. Thus from the above working we have

$$\tilde{S}(k,0) \sim \frac{|k|}{\pi \beta} + \frac{1}{2\pi \rho} \left( \frac{1}{\pi \beta} \right) \left( \frac{\beta - 2}{\beta} \right) k^2.$$  \hspace{1cm} (3.15)

This is precisely the result (2.57) expanded to second order in $k$ with $\beta' = \beta$, the substitution (2.45) and the partial derivative evaluated from the equilibrium equation of state $\beta_p = (1 - \beta/2)\rho$.

4. Conclusion

We have presented a systematic study of exact and asymptotic formulas relating to dynamical correlations for overdamped Brownian motion. Our main new results are the operator formula (2.19) for the current in the Fokker–Planck description, the generalized $f$-sum rule (2.28), the formula (2.35) relating the static current–current correlation to the static density–density correlation and the asymptotic formula (2.52) for the small-$k$ behaviour of the dynamical structure factor. We have illustrated these formulas using known exact evaluations of the dynamical density–density in certain special models.

Our work builds on the results presented in Refs. [1–3]. In this regard we point out that in Ref. [1] the current is defined not in terms of $dx_j(\tau)/d\tau$ as in (2.11), but rather with this factor replaced by $dx_j(\tau)/d\tau^{1/2}$. The latter choice is appropriate in applications of the Dyson Brownian motion model to chaotic spectra as the perturbing parameter $X$ is related to $\tau$ by $\tau = X^2$.

Acknowledgements

The visit of PJF to Orsay was financed by the Bede Morris fellowship scheme through his selection as The 1996 French Embassy Fellow.

Appendix

Here we will derive (2.57) using a thermodynamic argument. For notational convenience we will assume the domain is an interval of length $L$, although our argument applies equally well in higher dimensions. The setting is a uniform one-component system with a pair potential non-integrable at infinity. Such a system requires a neutralizing background for thermodynamic stability, and supposing the particles carry unit charge, the particle and charge density fluctuations are identical. We consider the microscopic density and suppose it has a fluctuation of the form

$$\delta \rho(x) = \rho_k e^{ikx} + \rho_k^* e^{-ikx}. \hspace{1cm} (A.1)$$
Here \( \rho_k \) is \( 1/L \) times the Fourier transform of the microscopic density. If \( 2\pi/k \) is a macroscopic wavelength, \( \rho_k \) is a good collective variable which can be used instead of the particle coordinates. The system will gain an electrostatic energy

\[
\frac{1}{2} \int_{-L/2}^{L/2} dx' \delta \rho(x') \int_{-L/2}^{L/2} dx \delta \rho(x)V_2(|x - x'|).
\]  
(A.2)

Let \( f(\rho) \) be the free energy per unit length of the locally neutral system. To second order in \( \delta \rho(x) \) the density fluctuation changes this free energy by an amount

\[
\delta f = \frac{\partial f}{\partial \rho} \bigg|_{\rho = \rho_0} \delta \rho(x) + \frac{1}{2} \frac{\partial^2 f}{\partial \rho^2} \bigg|_{\rho = \rho_0} (\delta \rho(x))^2.
\]  
(A.3)

Hence the total change in free energy due to the fluctuation is

\[
\delta F = \frac{1}{2} \int_{-L/2}^{L/2} dx' \delta \rho(x') \int_{-L/2}^{L/2} dx \delta \rho(x)V_2(|x - x'|) + \frac{1}{2} \frac{\partial^2 f}{\partial \rho^2} \bigg|_{\rho = \rho_0} \int_{-L/2}^{L/2} dx (\delta \rho(x))^2
\]

\[
\sim \frac{L|\rho_k|^2}{2} \left( \tilde{V}_2(k) + \frac{1}{2} \frac{\partial^2 f}{\partial \rho^2} \bigg|_{\rho = \rho_0} \right).
\]  
(A.4)

In general the probability density function for an event in the system is proportional to \( e^{-\beta \delta F} \) where \( \delta F \) is the corresponding change in the free energy. From (2.55), where now \( \delta \rho(k) = L \rho_k \), we have

\[
\tilde{S}(k) = L \langle |\rho_k|^2 \rangle.
\]  
(A.5)

Computing the r.h.s. of (A.5) from the Gaussian distribution \( e^{-\beta \delta F} \) gives (2.57).

References